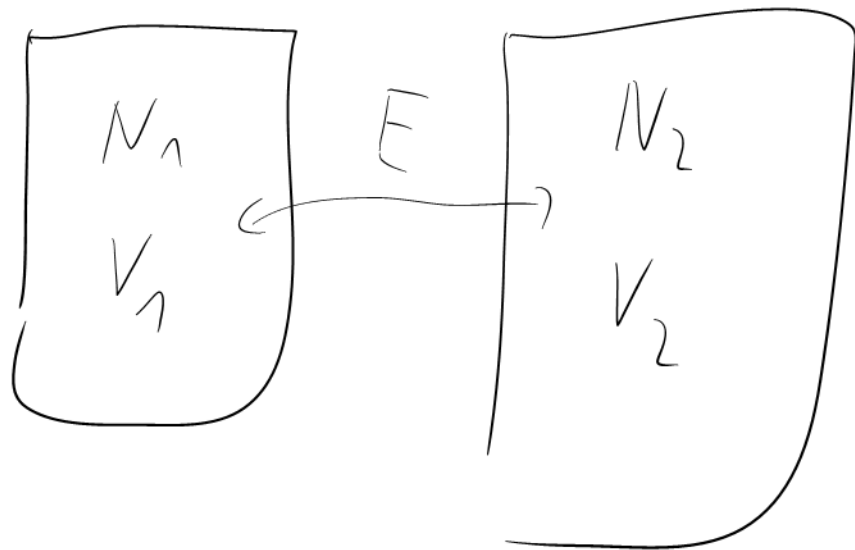


Boltzmann Distribution, Partition Function

and Free Energy



we know

$$\Omega_{\text{tot}}(E) = \frac{1}{\Delta E} \int_0^E dE_1$$

$$\Omega_1(E_1) \Omega_2(E - E_1) =$$

$$= \frac{\Delta E}{N_1! N_2!} \frac{1}{(2\pi\hbar)^{3(N_1+N_2)}} \int d\Gamma_1 \int d\Gamma_2$$

$$\delta(H_1(\Gamma_1) + H_2(\Gamma_2) - E)$$

we want: $p(\Gamma_1) d\Gamma_1 \equiv$ probability to find the

left system near phase space point $\Gamma_1 = ?$

we know:

$$p(\Gamma_1, \Gamma_2) = \frac{1}{\Omega_{\text{tot}}(E)} \frac{1}{N_1!} \frac{1}{N_2!} \frac{1}{(2\pi\hbar)^{3(N_1+N_2)}} \Delta E \int \delta(H_1(\Gamma_1) + H_2(\Gamma_2) - E)$$

$$p(\Gamma_1) = \int d\Gamma_2 p(\Gamma_1, \Gamma_2) = \frac{1}{\Omega_{\text{tot}}(E)} \frac{1}{N_1!} \frac{1}{(2\pi\hbar)^{3N_1}} \times$$

$$\times \frac{1}{N_2!} \frac{\Delta E}{(2\pi\hbar)^{3N_2}} \int d\Gamma_2 \int \delta(H_2(\Gamma_2) - (E - H_1(\Gamma_1))) =$$

$$p(\Gamma_n) = \frac{1}{N_n!} \frac{1}{(2\pi\hbar)^{3N_n}} \frac{\Omega_2(E - H_n(\Gamma_n))}{\Omega_{\text{tot}}(E)}$$

Now, $\Omega_2(E - H_n(\Gamma_n)) =$

$$= \exp\left[S(E - H_n(\Gamma_n))\right]$$

Now, let $N_2 \rightarrow \infty$, $V_2 \rightarrow \infty$, $\frac{N_2}{V_2}$ fixed

$E_2 \rightarrow \infty$, $\frac{E_2}{V_2}$ fixed \rightarrow infinite reservoir of energy
HEAT BATH

ONLY dependence
on Γ_n !

$$S(E - H_1(\Gamma_1)) = \underbrace{S(E)}_{O(N_2)} - \underbrace{\frac{\partial S}{\partial E}}_{= \beta = \frac{1}{T}} \underbrace{H_1(\Gamma_1)}_{O(N_1)} + \dots$$

neglect!

$$N_1 \ll N_2 \quad \frac{N_1}{N_2} \ll 1$$

→ Taylor expansion in $\frac{N_1}{N_2}$

$$\exp[S(E - H_1(\Gamma_1))] \simeq \exp(S(E)) \underbrace{\exp(-\beta H_1(\Gamma_1))}_{\text{Boltzmann factor}}$$

$$p(\Gamma_n) \propto \exp(-\beta H_n(\Gamma_n)) \quad \text{normalization}$$

$$\Rightarrow p(\Gamma_n) = \frac{1}{Z} \frac{1}{N_n!} \frac{1}{(2\pi h)^{3N_n}} \exp(-\beta H_n(\Gamma_n))$$

$$Z = \frac{1}{N_n!} \frac{1}{(2\pi h)^{3N_n}} \int d\Gamma_n \exp(-\beta H_n(\Gamma_n))$$

canonical partition function

Boltzmann

distribution

Remark:

$$Z(\beta) = \frac{1}{N_n!} \frac{1}{(2\pi h)^{3N_n}} \int dE_n \int d\Gamma_n \delta(U_n(\Gamma_n) - E_n)$$

$$\exp(-\beta E_n) =$$

$$= \int dE_n \Omega_n(E_n) \exp(-\beta E_n) \Rightarrow$$

$$\Omega_n(E_n) \xrightarrow[\text{transformation}]{\text{Laplace-}}, Z(\beta)$$

microcan.
part. fun.

canonical
part. fun.

In terms of potentials

$$dE = T dS - P dV + \mu dN \quad | \cdot \beta$$

$$F = E - TS \quad | \cdot (-\beta)$$

$$dF = -S dT - P dV + \mu dN$$

$$dS = \beta dE + \beta P dV - \beta \mu dN \quad S$$

$$-\beta F = S - \beta E$$

$$d(-\beta F) = -E d\beta + \beta P dV - \beta \mu dN$$

and

$$\frac{\partial(-\beta F)}{\partial \beta} = -E$$

↓ Legendre tr.

$-\beta F$

on the other hand:

$$\langle E \rangle = \frac{\int d\Gamma \exp(-\beta H(\Gamma)) (-H(\Gamma))}{\int d\Gamma \exp(-\beta H(\Gamma))} =$$

$$= \frac{\frac{\partial}{\partial \beta} \int d\Gamma \exp(-\beta H(\Gamma)) \frac{1}{N!} \frac{1}{(2\pi h)^{3N}}}{\int d\Gamma \exp(-\beta H(\Gamma)) \frac{1}{N!} \frac{1}{(2\pi h)^{3N}}} =$$

$$= \frac{1}{Z} \frac{\partial}{\partial \beta} Z = \frac{\partial}{\partial \beta} \ln Z$$

$$\text{So: } -\bar{E} = \frac{\partial}{\partial \beta} (-\beta \bar{F}) = \frac{\partial}{\partial \beta} \ln Z$$

$$\rightarrow \text{SUSPICION: } -\beta \bar{F} = \ln Z$$

What about the V derivative? We know

$$\frac{\partial}{\partial V} (-\beta \bar{F}) = \beta P, \quad \text{on the other hand}$$

$$\frac{\partial}{\partial V} (\ln Z) = \frac{1}{Z} \frac{\partial Z}{\partial V} = \frac{1}{Z} \frac{1}{N!} \frac{1}{(2\pi h)^{3N}} \int d\mathbf{r} e^{-\beta H(\mathbf{r})} (-\beta) \frac{\partial H}{\partial V} =$$

$$-\beta \left\langle \frac{\partial H}{\partial V} \right\rangle_{\text{canon.}} = -\beta \left\langle \frac{\partial H}{\partial V} \right\rangle_{\text{microcan.}}$$

thermodyn.
limit

$$= -\beta \left. \frac{\partial E}{\partial V} \right|_S = (-\beta)(-P) = \beta P$$

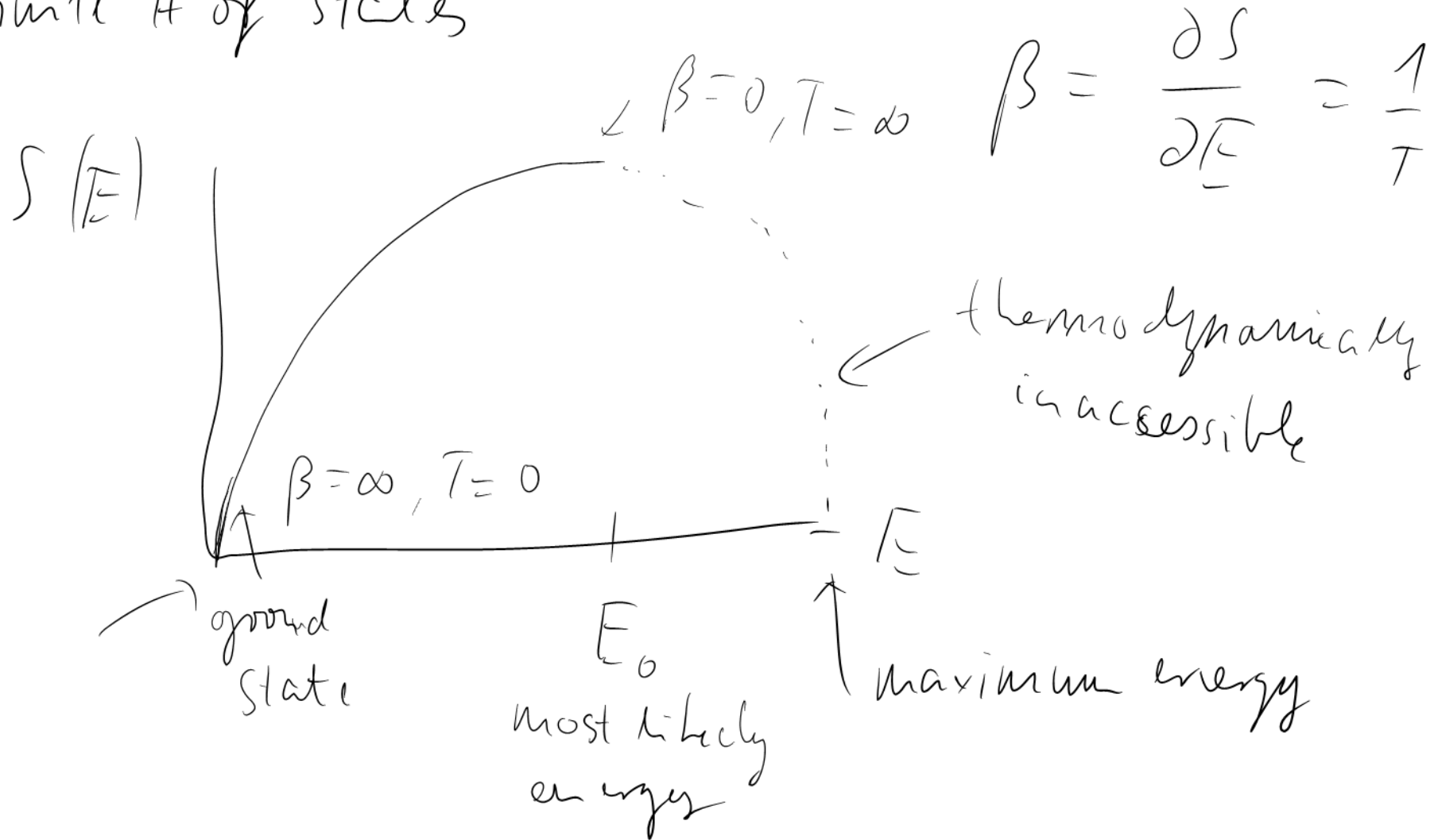
$$dE = T dS - P dV + \mu dN$$

$$\Rightarrow \left(\frac{\partial}{\partial V} (-\beta F) = \frac{\partial}{\partial V} \ln Z \right)$$

$$\Rightarrow -\beta F = \ln Z + \underbrace{\text{const.}}_{= 0 \text{ why?}}$$

Show that the constant is zero

Special system: Energy is bounded, system has a finite # of states



$$\Omega(E_0) = e^{S(E_0)} \simeq \# \text{ of ALL states in the system}$$

$$\beta = 0 \rightarrow \bar{E} = E_0, \quad S(\beta = 0) = \ln \Omega(E_0) \\ \simeq \ln(\# \text{ of all states})$$

on the other hand:

$$-\beta F = S - \beta \bar{E} \xrightarrow{\beta \rightarrow 0}$$

$$\rightarrow S(\beta = 0) - 0 \cdot E_0 = S(\beta = 0) \\ = \ln(\# \text{ of all states})$$

finally:

$$\ln Z(\beta \rightarrow 0) = \ln \underbrace{\frac{1}{N!} \frac{1}{(2\pi h)^{3N}} \int d\Gamma 1}_{\# \text{ of all states}} =$$

$$= \ln (\# \text{ of all states}) = S(\beta = 0)$$

$$\Rightarrow \text{const.} = 0, \quad -\beta F = \ln Z$$

$$\boxed{F = -T \ln Z}$$

□
0

The Canonical Partition Function

Species $Z = \frac{1}{N!} \frac{1}{(h^3)^N} \int d\Gamma e^{-\beta H}$ Cartesian coordinates

$$d\Gamma = d^3\vec{p}_1 d^3\vec{p}_2 \dots d^3\vec{p}_N d^3\vec{r}_1 \dots d^3\vec{r}_N$$

$$H = \sum_i \frac{1}{2m_i} \vec{p}_i^2 + U(\{\vec{r}_i\})$$

$d^3\vec{p}_i = dp_{ix} dp_{iy} dp_{iz}$

$$e^{-\beta H} = \exp\left(-\beta \frac{p_{1x}^2}{2m_1}\right) \exp\left(-\beta \frac{p_{1y}^2}{2m_1}\right) \dots \exp\left(-\beta \frac{p_{Nx}^2}{2m_N}\right) e^{-\beta U}$$

Gaussian integrals \Rightarrow

$$Z = \frac{1}{(2\pi\hbar)^{3N}} \quad Z_{\text{momenta}} \quad Z_{\text{coord.}}$$

$$Z_{\text{coord.}} = \left(\prod_{i=1}^N \frac{1}{h^{3_i}} \right)$$

$$\int d^3\vec{r}_1 \dots \int d^3\vec{r}_N e^{-\beta U}$$

$$Z_{\text{momenta}} = \int d^3\vec{p}_1 e^{-\beta \frac{\vec{p}_1^2}{2m_1}} \dots \int d^3\vec{p}_N e^{-\beta \frac{\vec{p}_N^2}{2m_N}}$$

$$= \underbrace{\sqrt{2\pi m_1 T}}^3 \dots \underbrace{\sqrt{2\pi m_N T}}^3 \quad (3N \text{ roots})$$

Define:

$$\frac{\sqrt{2\pi m_i T}}{2\pi\hbar} =: \frac{1}{\lambda_i}$$

λ_i : "thermal de Broglie wavelength" of particle i

$$Z = Z_{\text{coord}} \cdot \prod_i \frac{1}{\lambda_i^3}$$

for ideal gas: $Z = \left(\prod_k \frac{1}{N_k!} \right) \left(\prod_i \frac{1}{\lambda_i^3} \right) V^N$

$$\rightarrow \bar{F} = -T \ln Z = -TN \ln V + \text{terms indep. of } V$$

$$P = - \left. \frac{\partial \bar{F}}{\partial V} \right|_T = + TN \frac{1}{V} \Rightarrow \boxed{PV = TN}$$

probability for the velocity of particle i :

$$P(\vec{v}_i) = \frac{1}{\sqrt{2\pi \frac{T}{m_i}}^3} \exp\left[-\beta \frac{m_i}{2} \vec{v}_i^2\right]$$

Maxwell velocity distribution

$$\langle v_{ix}^2 \rangle = \langle v_{iy}^2 \rangle = \langle v_{iz}^2 \rangle = \frac{T}{m_i}$$

$$\frac{m_i}{2} \langle v_{ix}^2 \rangle = \frac{T}{2} \quad \text{equipartition theorem:}$$

$T/2$ per "quadratic" degree of freedom